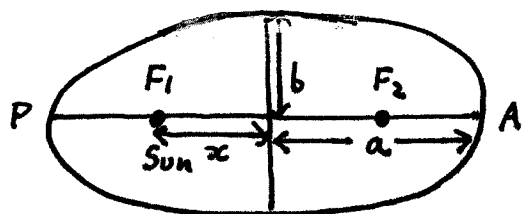


## The basic ellipse

The size and shape of the ellipse are defined in terms of its semi-major and semi-minor axes,  $a$  and  $b$ , and its eccentricity,  $e$ . These quantities are related:

$$\frac{b^2}{a^2} = 1 - e^2$$

If  $e = 0$ , the figure is a circle ( $a = b$ ).  
 (If  $e = 1$ , the figure becomes a parabola and if  $e > 1$ , the figure is a hyperbola. The distances,  $x$ , of the two foci,  $F_1$  and  $F_2$ , from the centre of the ellipse are given by the product of the semi-major axis and the eccentricity, that is



$$x = \pm a e$$

$A =$  aphelion  
 $P =$  perihelion

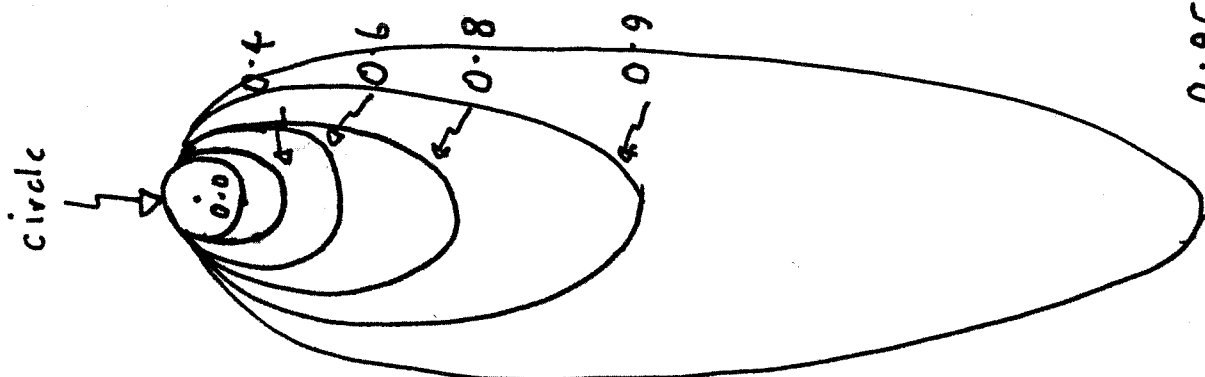
It is useful to be able to calculate the distances from the Sun of the perihelion and aphelion points,  $q$  and  $Q$ , respectively. These quantities are related to  $a$  and  $e$  by the following expressions:

$$q = a(1 - e)$$

$$Q = a(1 + e)$$

As  $e$  has no units associated with it, the equations "work" regardless of the units of  $q$  and  $Q$  (as long as they

are all in the same units as each another). If the value of the perihelion distance is held constant and the eccentricity increased, then the semi-major axis (and thus the aphelion distance) must also increase.



This rough sketch shows the general idea.

DF  
 2015, Sept. 10